

Geometry for Quantum Science Group  
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# Quantisation and Path Integrals

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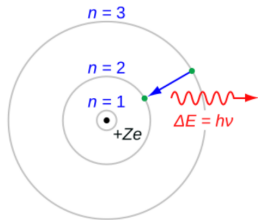


- 1 The end of the classical world
- 2 Methods of quantisation
- 3 Feynman's path integral
- 4 The relativistic path integral

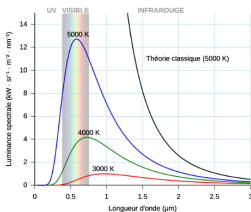


# The dark cloud above classical physics

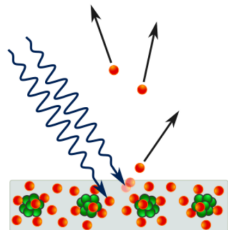
- Atomic spectra
- Black body radiation
- Photoelectric effect



(a) JabberWork, File :Bohr atom model.svg



(b) Darth Kule, File :Black body-fr.svg



(c) Ponor, File :Photoelectric effect in a solid - diagram.svg

→ from continuous to discrete observables.



# The Bohr-Sommerfeld rule

*Quantum numbers*: adiabatic invariants of quantised systems.

## Bohr-Sommerfeld rule

Let  $q$  be a periodic coordinate with conjugate momentum  $p$ . Then there is a quantum number  $n$  such that

$$\oint p \, dq = nh \quad (1)$$

where  $h$  is Planck's constant (the quantum of action).

→ from dynamics to statics.



[picture of phase space]

Observable in classical physics: function on phase space.

Quantum observable: integral of a classical observable / function of phase space curves.



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The five postulates:

- ① Superposition principle
- ② Quantisation principle
- ③ Born's rule
- ④ Spontaneous collapse
- ⑤ Schrödinger's equation

How do we make sense of 2.?

→ connecting classical to quantum observables.

**Formally replace** in the Hamiltonian all observables by operators.



Since classical observables are functions of  $(q_k, p_k)$ , quantum observables become functions of  $(\hat{q}_k, \hat{p}_k)$  (forgetting about spin).

→ just describe the algebra of canonical variable operators:

## Heisenberg commutation relations

$$[\hat{q}_i, \hat{p}_j] = i\hbar\delta_{ij} \quad (2)$$

$$[\hat{q}_i, \hat{q}_j] = 0 \quad (3)$$

$$[\hat{p}_i, \hat{p}_j] = 0 \quad (4)$$

→ enforces Heisenberg indetermination relations.





More generally:

$$[\hat{f}, \hat{g}] = i\hbar \widehat{\{f, g\}} \quad (5)$$

But...

## Groenewold theorem

There is no map from phase space to the space of operators that simultaneously

- ① sends 1 onto  $\hat{\mathbb{I}}$ ;
- ② sends  $q_k$  onto  $q_k \cdot$  and  $p_k$  onto  $i\hbar \frac{\partial \cdot}{\partial q_k}$ ;
- ③ preserves polynomials;
- ④ satisfies eqn 5.



Best we can do: only preserve polynomials up to degree 3.

→ **Weyl's transform** (tiens tiens).

Leads to **deformation quantization** when generalized to arbitrary Poisson manifolds.

→ Weyl's transform replaced by the Kontsevich quantisation formula.

Physical interpretation becomes awkward

+ Lorentz-covariance not explicit for relativistic theories

+ Heisenberg's indeterminacy: quantum states are not even measurable...

Should we do something *completely different*?



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# The sum-over-paths picture

From Bohr-Sommerfeld theory:

- quantum states are curves in phase space (paths in physical space);
- quantum observables are functions of phase space curves.

From Young's slits experiment:

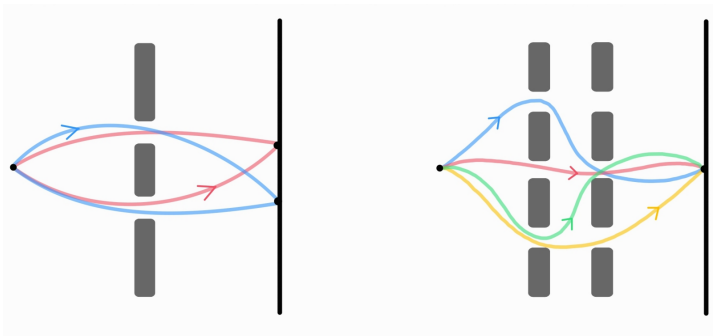
- a particle passes through both slits since both paths *interfere*;
- this generalizes to an arbitrary number of slits.

From the correspondence principle:

- in the classical limit, there is only one path that counts.



# The sum-over-paths picture



What is the probability to propagate from  $x_0$  to  $y_1$ ? to  $y_2$ ?

$$G(x_0, t_0, x_1, t_1) = \langle x_1 | e^{-i(t_1-t_0)\hat{H}/\hbar} | x_0 \rangle = \sum_n \langle x_1 | E_n \rangle \langle E_n | e^{-i(t_1-t_0)\hat{H}/\hbar} | x_0 \rangle \quad (6)$$

$$= \sum_n \psi_n(x_1) \psi_n^*(x_0) e^{-i(t_1-t_0)E_n/\hbar} \quad (7)$$



# Definition of the path integral

$$G(x_0, t_0, x_1, t_1) = \int_{x(t_0)=x_0}^{x(t_1)=x_1} \mathcal{D}x(t) e^{iS[x]/\hbar} \quad (8)$$

where we apply the **time-slicing** procedure

$$\mathcal{D}x(t) = \lim_{N \rightarrow \infty} e^{-i\pi/4} \sqrt{\frac{m}{2\pi\hbar\epsilon}} \prod_{n=1}^{N-1} \left[ \int dx_n e^{-i\pi/4} \sqrt{\frac{m}{2\pi\hbar\epsilon}} \right] \quad (9)$$

with  $N = (t_1 - t_0)/\epsilon$

→ classical limit.



# Propagators and Green functions

$G(x_0, t_0, x_1, t_1)$  is called the **propagator** of the particle.

Where does it come from?

$$i\hbar \frac{\partial}{\partial t} G = H G \quad (10)$$

$$G(x_0, t_0, x_1, t_0) = \delta(x_1 - x_0) \quad (11)$$

and defining the *retarded propagator*  $G_R = G\Theta(t_1 - t_0)$

$$\left( i\hbar \frac{\partial}{\partial t} - H \right) G_R = \delta(x_1 - x_0) \delta(t_1 - t_0) \quad (12)$$

$$G_R(x_0, t_0, x_1, t_0) = \delta(x_1 - x_0) \quad (13)$$

→  $G_R$  is the **Green function** associated to Schrödinger's equation.



Let  $D$  be a linear differential operator. Solve for  $\phi$  the partial differential equation

$$D\phi(x) = j(x). \quad (14)$$

A **Green function** for  $D$  is any solution of

$$DG(x_0, x) = \delta(x - x_0). \quad (15)$$

## Green functions theorem

Let  $G$  be a Green function for  $D$ . Then

$$\phi(x) = \int G(x_0, x)j(x_0) dx_0 \quad (16)$$

is a solution of eqn 14.





Nonrelativistic quantum mechanics = 1D field theory

Generalization to arbitrary quantum field theories

$$G(\varphi, \mathcal{N}) = \int_{\phi|_{\partial\mathcal{N}=\varphi}} \mathcal{D}\phi(x) e^{iS[\phi]/\hbar} \quad (17)$$

with  $\mathcal{N}$  an arbitrary spacetime domain with  $\varphi$  a field on  $\partial\mathcal{N}$ .

→ explicit Lorentz-covariance.

However, quantization and unitarity are not explicit anymore.



Time  $\longrightarrow$  Temperature.

## Partition function

$$Z(\beta) = \text{Tr} e^{-\beta\hat{H}} = \int dx \langle x | e^{-\beta\hat{H}} | x \rangle = \oint_{t_1-t_0=-i\hbar\beta} \mathcal{D}x(t) e^{iS[x]/\hbar} \quad (18)$$

then

$$\langle \hat{O} \rangle_\beta = \frac{1}{Z(\beta)} \int dx \langle x | e^{-\beta\hat{H}} \hat{O} | x \rangle = \frac{1}{Z(\beta)} \int_{t_1-t_0=-i\hbar\beta} dx dy G(y, t_0, x, t_1) \langle y | \hat{O} | x \rangle. \quad (19)$$



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# Large time limit and expectation values

$$G(x_0, t_0, x_1, t_1) = \sum_n \psi_n(x_1) \psi_n^*(x_0) e^{-i(t_1 - t_0) \hat{E}_n / \hbar}. \quad (20)$$

Setting  $E_0 = 0$  and taking  $t_1 \rightarrow -i\infty$  and  $t_0 \rightarrow +i\infty$  yields

$$G_\infty = \psi_0(x_1) \psi_0^*(x_0) = \langle x_1 | E_0 \rangle \langle E_0 | x_0 \rangle \quad (21)$$

→ probability amplitude for a particle in the ground state to be at  $x_0$  and  $x_1$   
→ *field* point of view.

$$\langle 0 | \hat{O} | 0 \rangle = \int_{\mathcal{M}} \mathcal{D}\phi O[\phi] e^{iS[\phi]/\hbar} \quad (22)$$

on the whole manifolds  $\mathcal{M}$  → time-ordered product.



## Chapman-Kolmogorov equation

$$G(x_0, t_0, x_1, t_1) = \int dx G(x_0, t_0, x, t) G(x, t, x_1, t_1) \quad (23)$$

## Sewing law

$$\int_{\Sigma_1 \rightarrow \Sigma_2} \mathcal{D}\phi e^{iS[\phi]/\hbar} = \int_{\Sigma'} \mathcal{D}\varphi' \int_{\Sigma_1 \rightarrow \Sigma_2}^{\phi|_{\Sigma'} = \varphi'} \mathcal{D}\phi e^{iS[\phi]/\hbar}. \quad (24)$$



Equal time canonical anticommutation relations

$$\left\{ \psi_\alpha(x), \psi_\beta^\dagger(y) \right\} = \delta^{\alpha\beta} \delta^{(3)}(x - y) \quad (25)$$

Grassmann variables: Grassman algebra of anticommuting generators  
 $\{\theta_1, \theta_2\}$

$$\theta_1 \theta_2 = -\theta_2 \theta_1 \quad (26)$$

→ each  $\psi(x)$  and  $\psi^\dagger(x)$  for each spacetime point are Grassmann variables.



Nonrelativistic path integral is defined through **Wiener measure**

→ connection with Schrödinger's equation ensured by the Feynman-Kac formula.

However

Ill defined in general...

→ perturbative quantum field theory

→ functorial field theory.



# Acknowledgements

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